

PLANTING COMPACT BINARIES IN COSMOLOGICAL SIMULATIONS

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NEW FIELD IN ASTRONOMY

Fundamental physics

Beyond GR
Mass of graviton
Hubble constant

Nuclear physics

Neutron star equation of state
R-process elements

Global star formation

Initial mass function
Low-metallicity environments
R-process elements
Dwarfs galaxies
Milky-Way structure

Gravitational waves

from stellar remnants
(black holes, neutron stars,
white dwarves)

High energy astrophysics

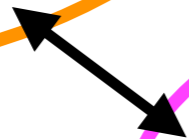
Relativistic outflows
BH formation
Gamma-ray bursts

Stellar/Binary evolution

Supernova kicks
Common envelope
Mass loss
Cluster evolution

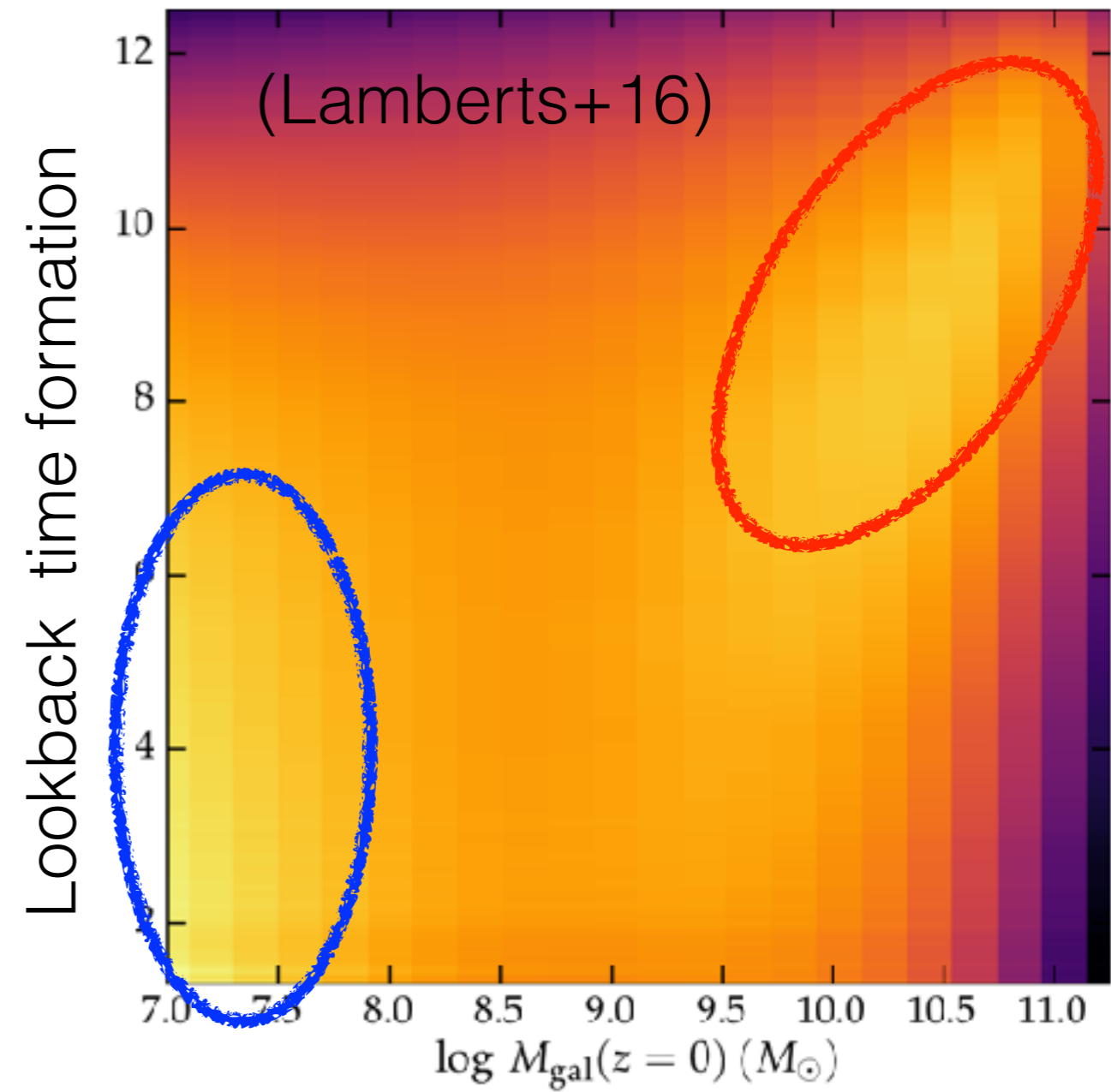
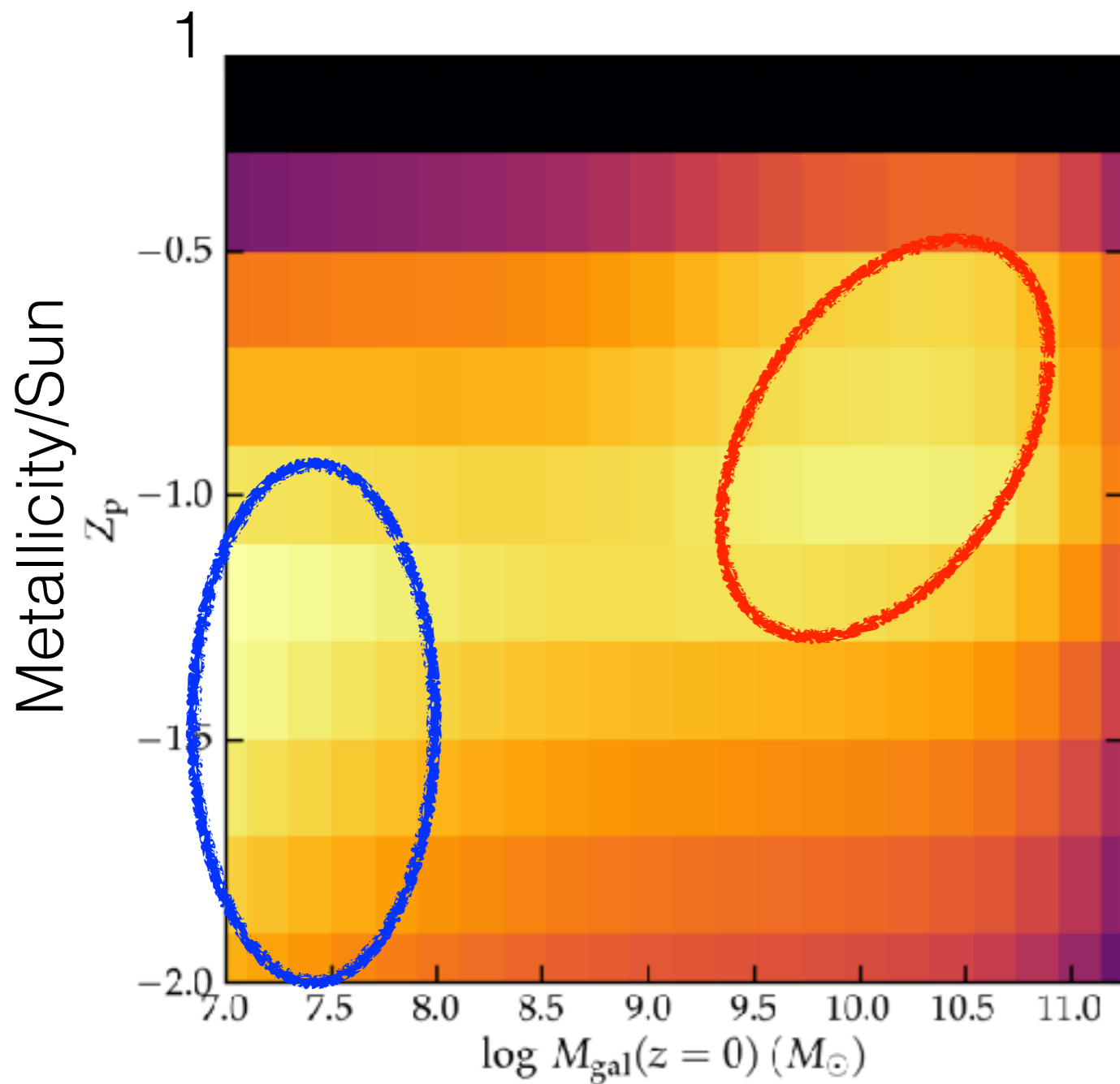
And numerics for

Waveforms
Data analysis



FORMATION OF GW PROGENITORS

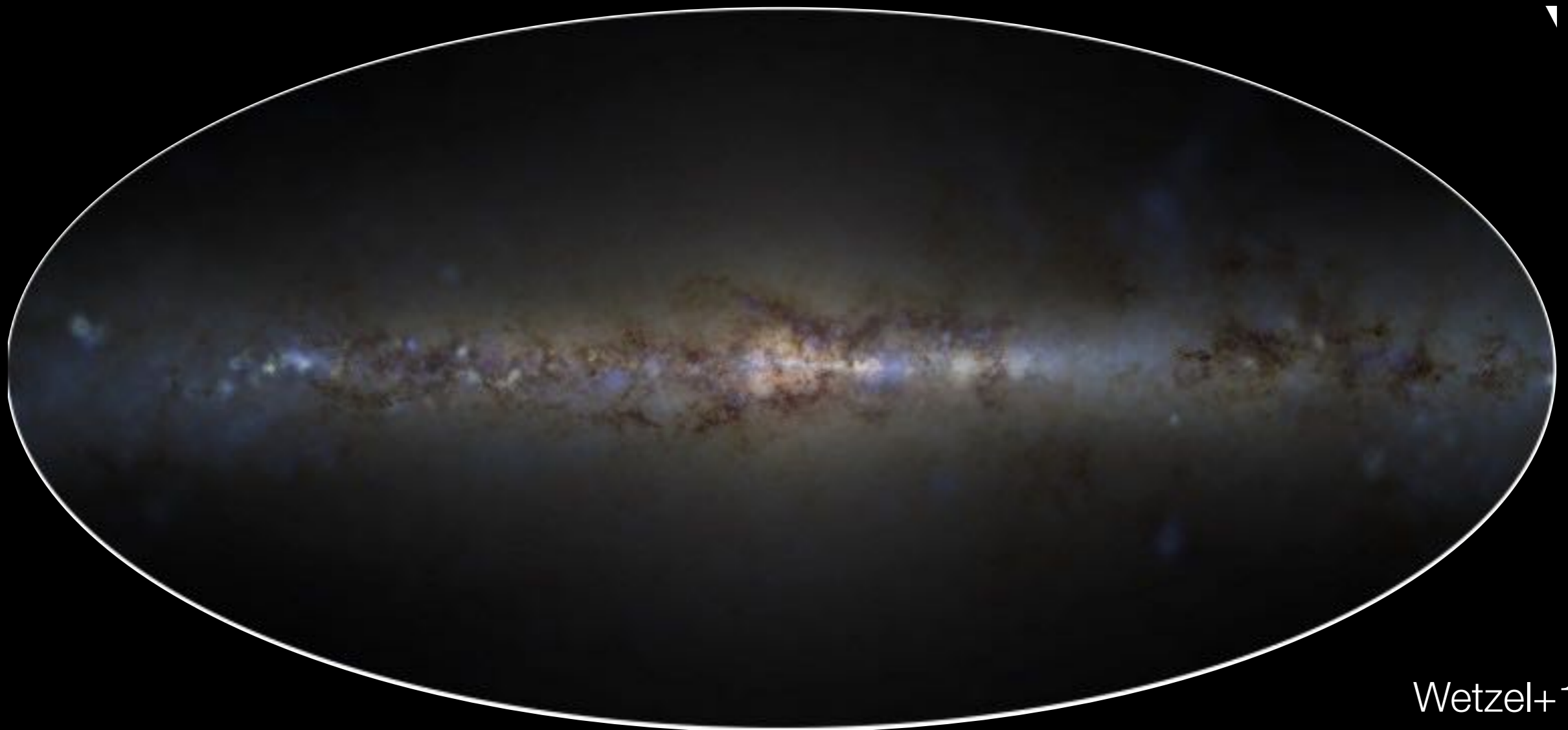
Dwarf galaxy Milky Way



COMBINE MERGER MODELS WITH COSMOLOGICAL SIMULATIONS

FIRE SIMULATIONS

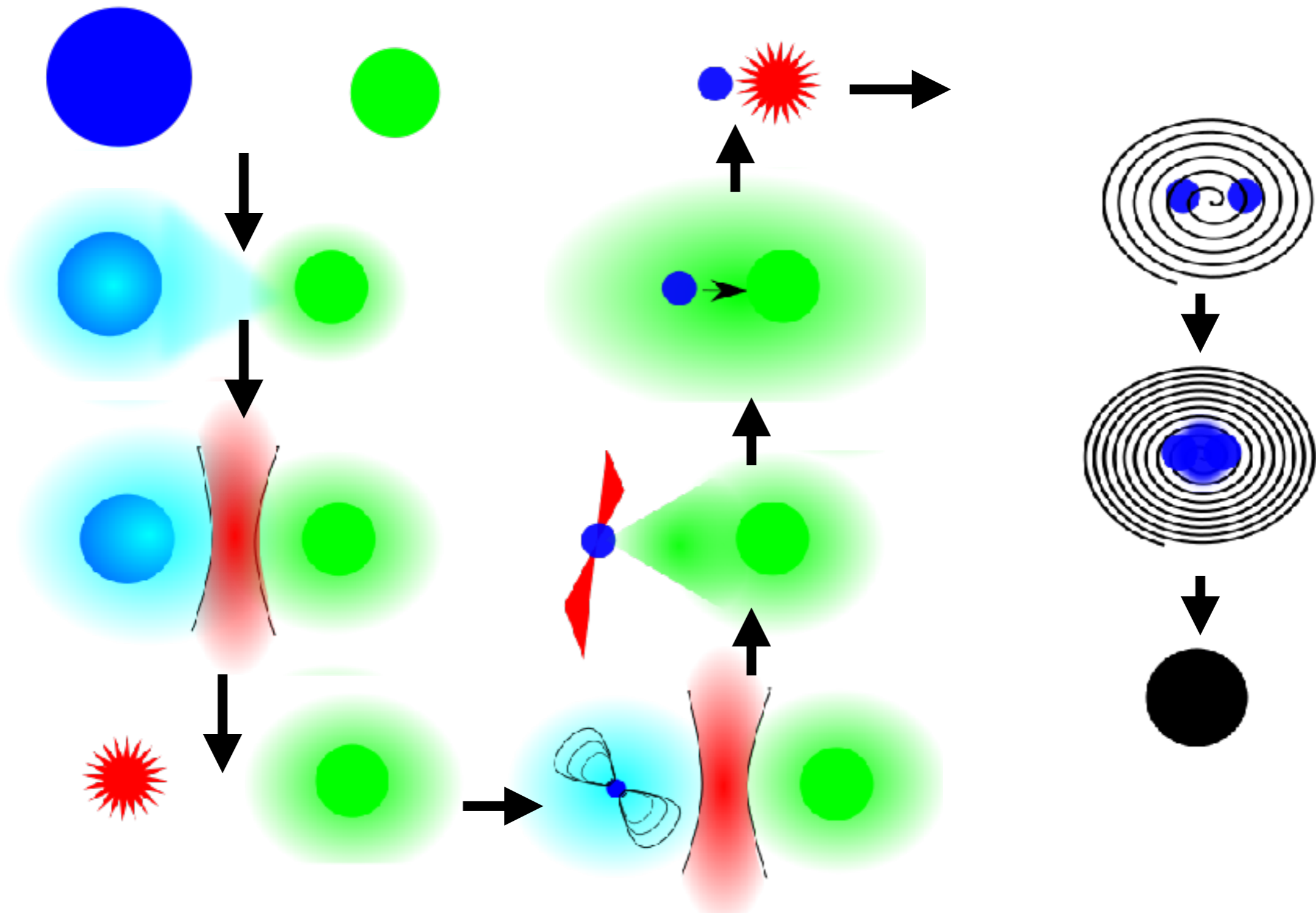
Feedback In Realistic Environments project (FIRE, Hopkins+14,17)



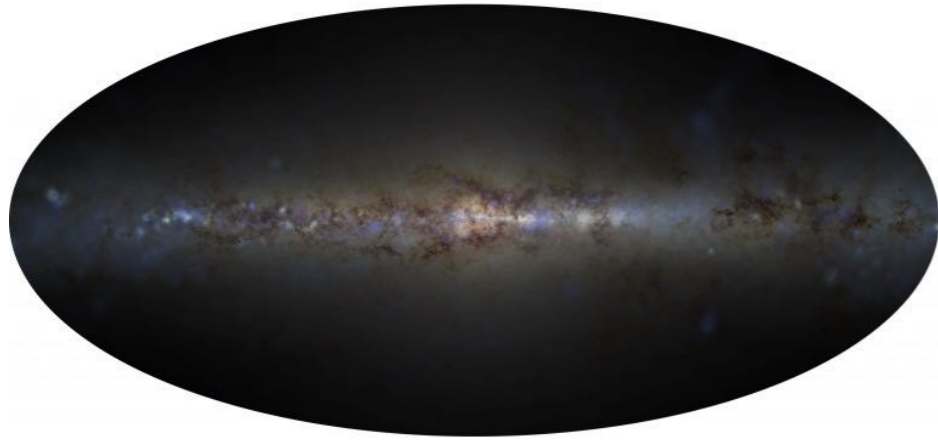
Wetzel+16

Ex: LATTE simulation Mass resolution : 7000 Msun

MAKING BINARY COMPACT OBJECTS

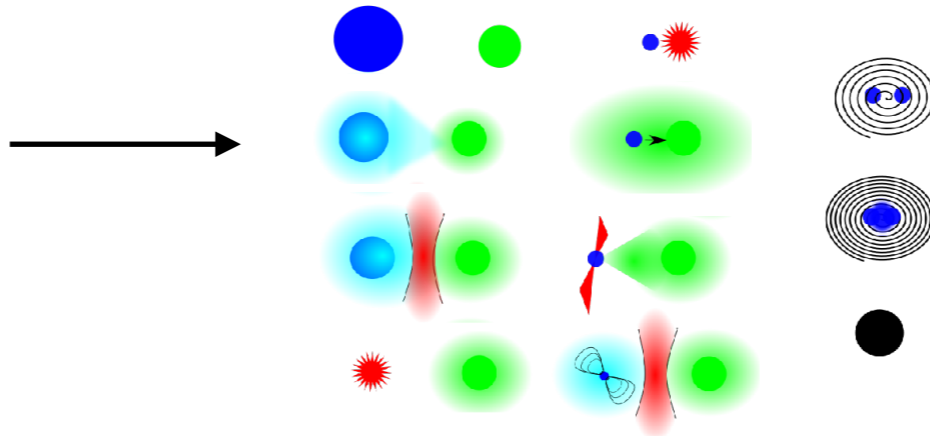


GROWING COMPACT BINARIES IN GALAXIES



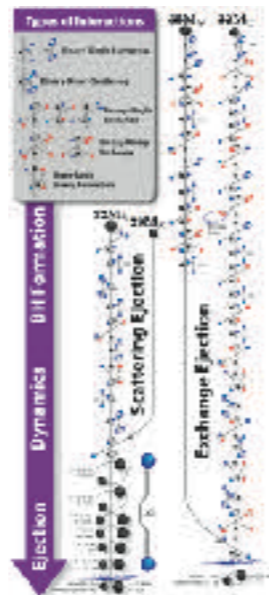
Star formation history
Metallicity
Positions/Trajectory
Gas properties

Star cluster properties
(M. Grudic)

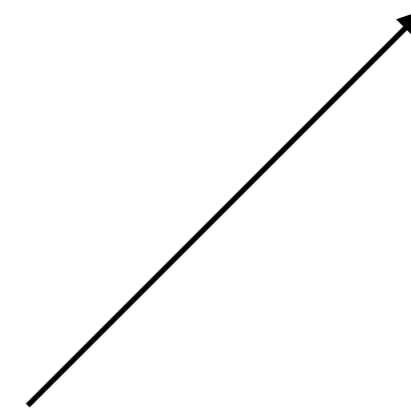


Binary population model
Metallicity dependent

Gravitational wave emission

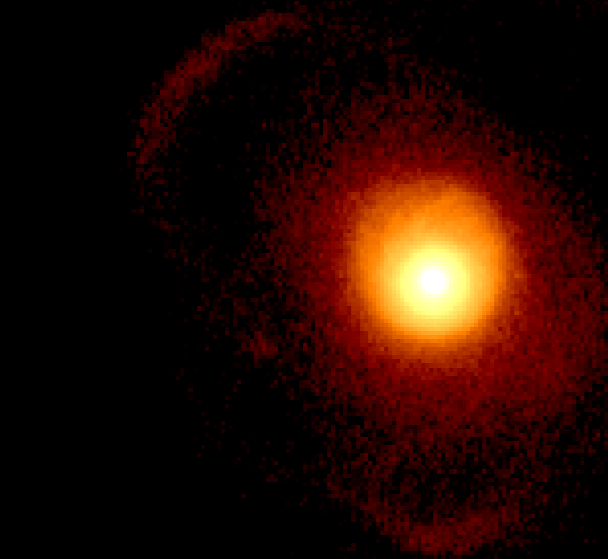


cluster mergers
(C. Rodriguez)

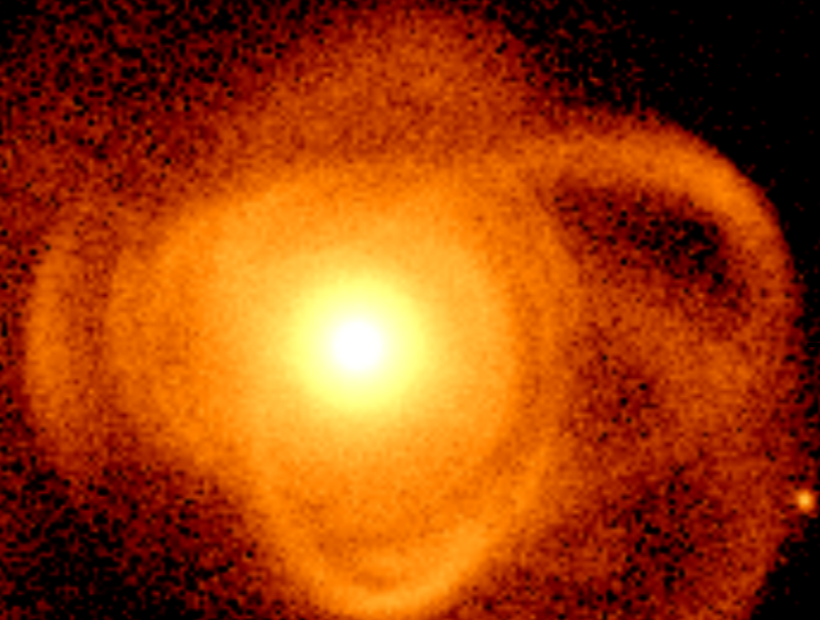


IMPACT OF COMPLEX STAR FORMATION

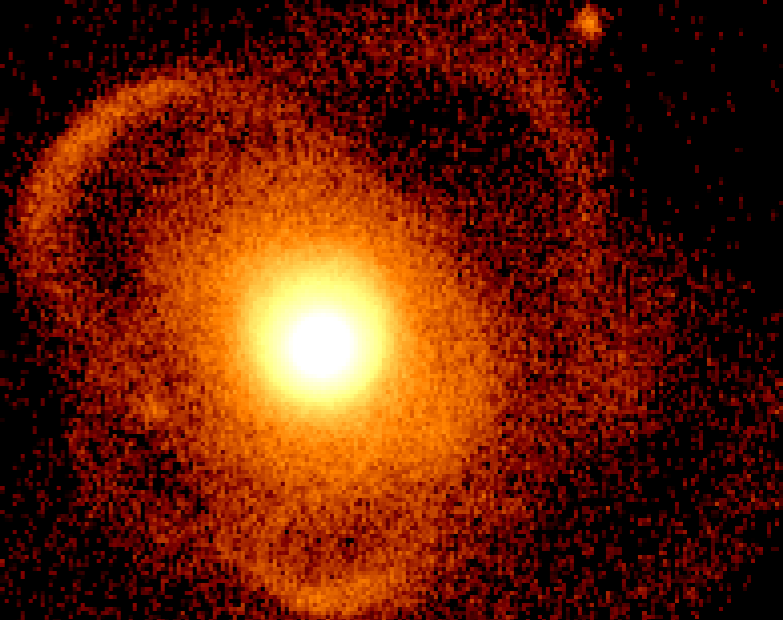
100 kpc
←→



$>10^{10}$ stars



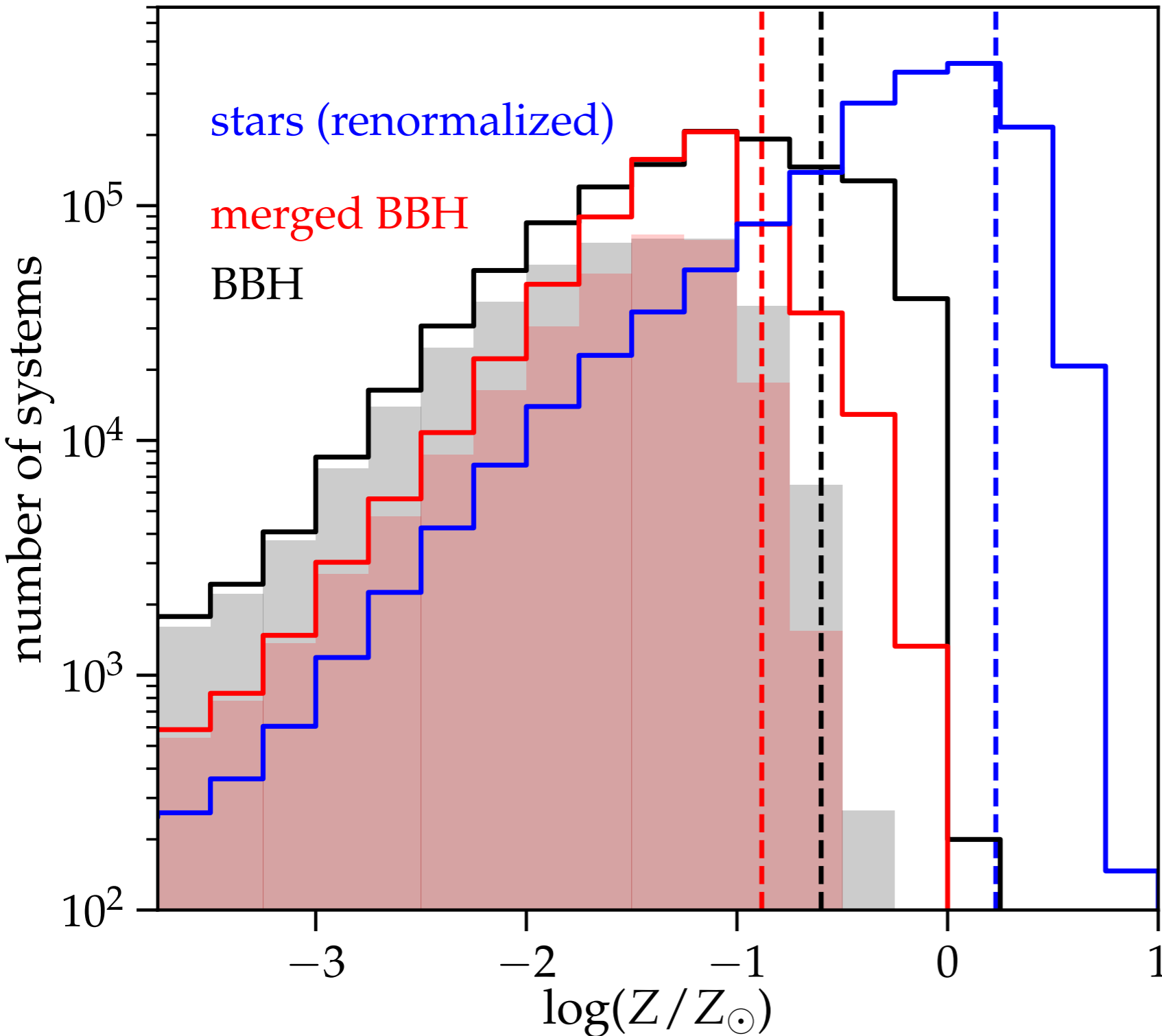
1 million binary
black holes



100 million binary
white dwarfs

Different populations found in different structures

IMPACT OF DWARF GALAXIES



Mean metallicity
~ 20% of Sun

1/3 of black holes were
Formed outside Milky Way
-> dwarf galaxies matter
-> needs high resolution

"INDIVIDUAL" NUMERICAL REQUIREMENTS

FIRE simulations : MW-like galaxies:

10 runs with 1-5 million CPU ($M \sim 7000 M_{\text{sun}}$)

1 run with 25 million CPU in progress ($M \sim 900 M_{\text{sun}}$)

Outputs : 7-20 GB/snapshot x 600 : 5-10 TB/run

National/European-scale supercomputers

Cluster evolution : 100 000 CPU hours, need ~ 20 models
scaling difficult beyond ~ 100 CPU, specific supercomputers

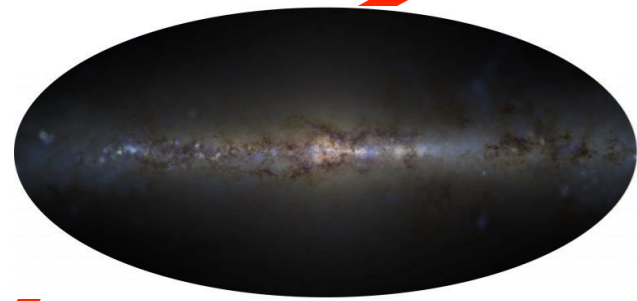
10 million binaries: ~ 24 CPU hours

Need ~ 100 models for white dwarfs

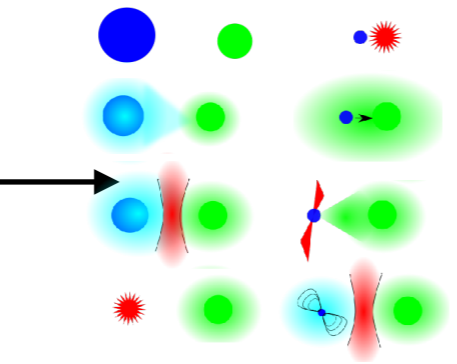
~ 700 models for black holes

Embarrassingly parallel, ok on local clusters, needs
"bookkeeping"

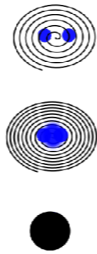
BRINGING IT ALL TOGETHER



Star formation
Metallicity



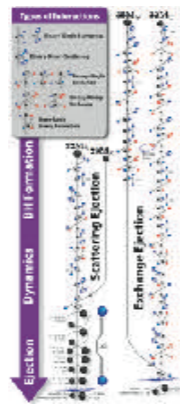
Binary model



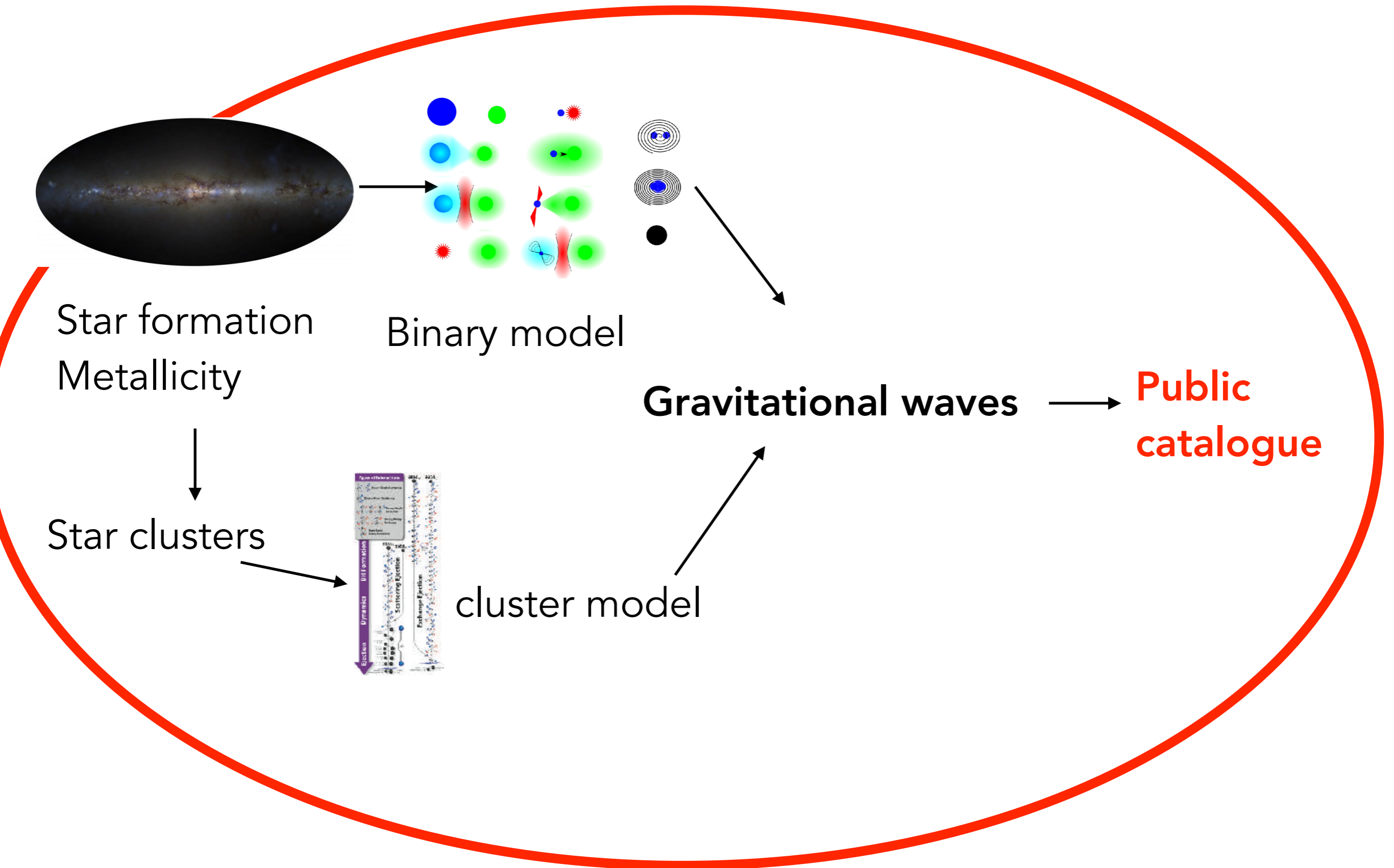
Gravitational waves

Public
catalogue

Star clusters



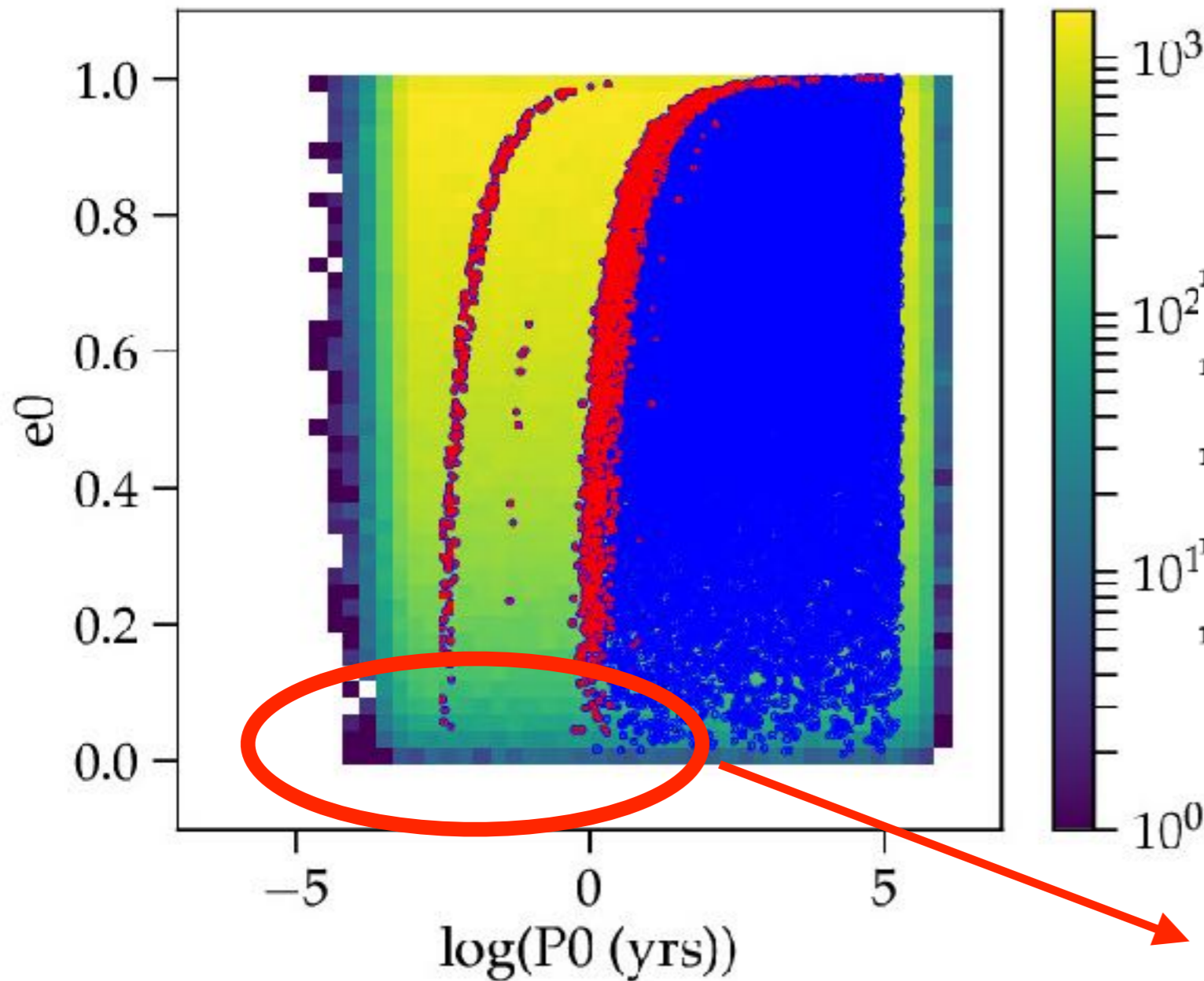
cluster model



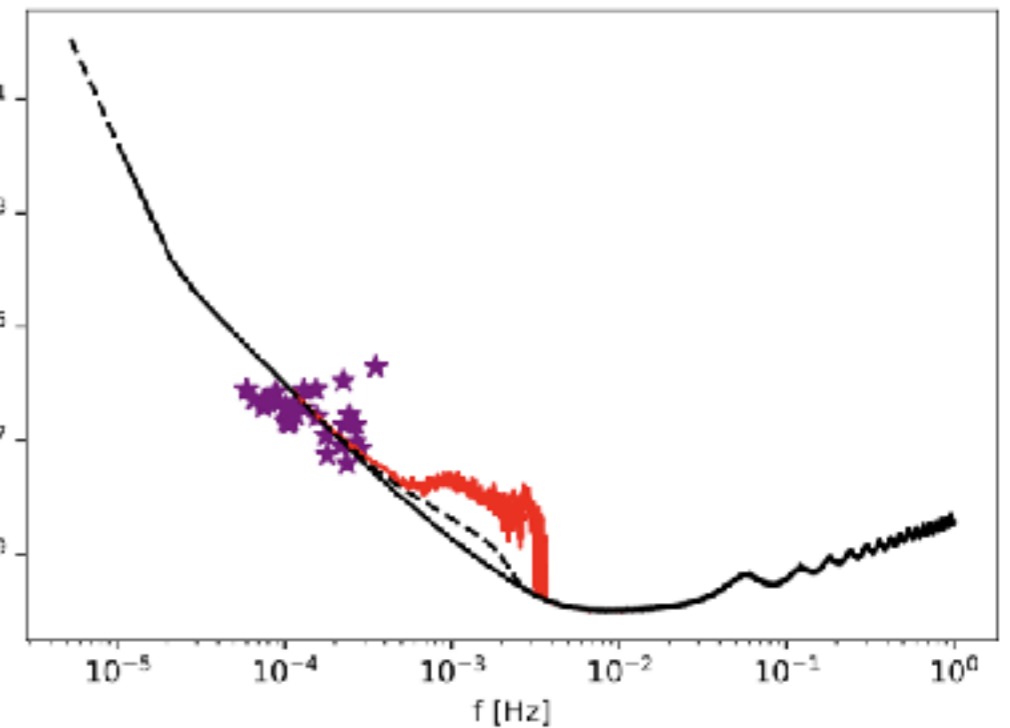
BRINGING IT ALL TOGETHER

1) MW simulations for LISA predictions

Sampling issues : initial conditions need to be sampled properly (x10)



Galactic background in LISA



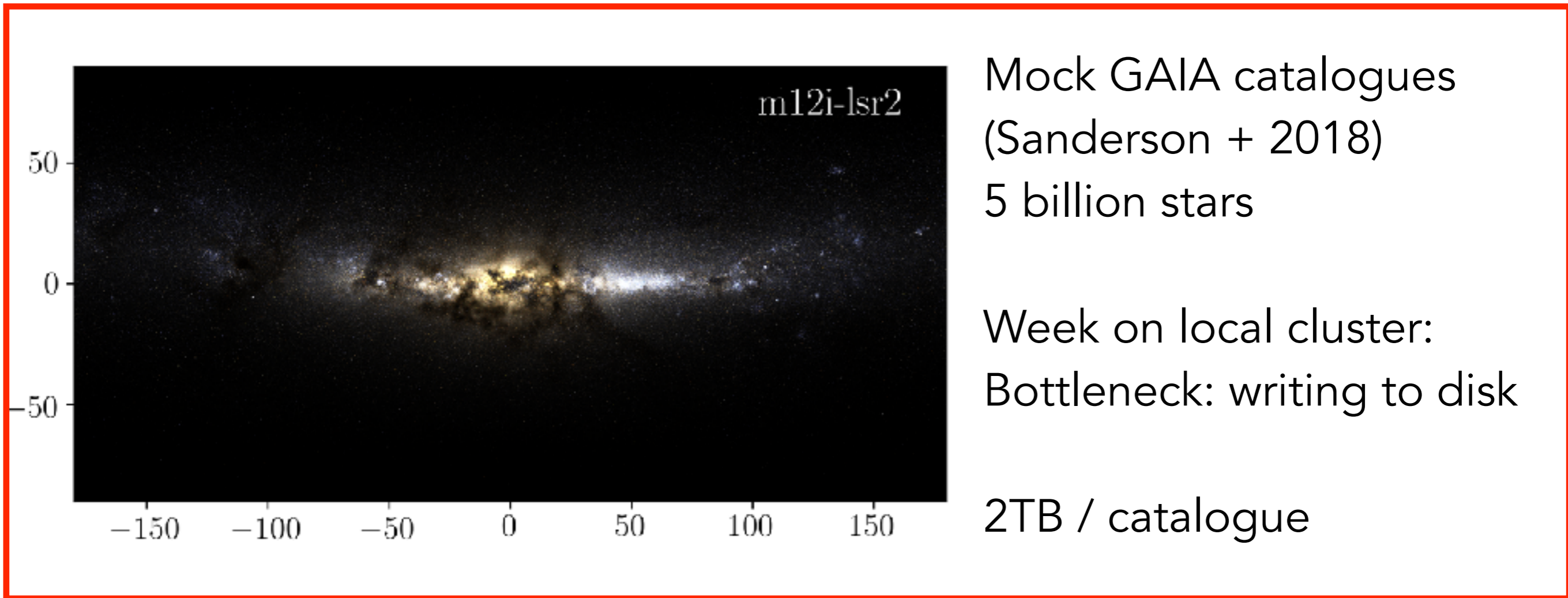
Missing binaries =
underestimated background

BRINGING IT ALL TOGETHER

2) Communicating results and accessibility

Making a catalogue: *accessible, durable, easy to use, advertised* : how?

2D/3D visualisation



In prep: double white dwarf catalogues (~100 million objects)

BRINGING IT ALL TOGETHER

3) LIGO/Virgo mergers from binaries and clusters

FIREBOX simulation. 15 Mpc/h, $M \sim 60\,000 M_{\text{sun}}$

~ 50 MW galaxies, 1000 dwarf galaxies

5 million CPU (PRACE, R. Feldmann), 50 TBs, 200 GB/output



Needs for analysis:

Storage and high-memory computing (finding cloud properties,

Determine galactic tides, star particle history)

Flexible outputs, good time sampling (for gas evolution)

Intermediate data products: which formats?

Visualisation: tools?

Statistical tools to analyze different models